

## Astronomy GR6001: Problem Set #4

Due in my office by Tuesday, November 2, 2021

### Problem 1 (25 points):

Consider the most general form of a single monochromatic wave propagating in the direction  $\hat{k}$  at frequency  $\omega$ , i.e. an elliptically polarized wave. Compute  $d\sigma/d\Omega$ , the differential cross-section for scattering such a wave off a free charge as a function of the direction towards the observer  $\hat{n}$  (i.e. the Thomson cross-section for an elliptically polarized wave).

### Problem 2 (25 points):

Consider two particles, with electric charges  $q_1$  and  $q_2$ , and masses  $m_1$  and  $m_2$ , orbiting at non-relativistic speeds in each other's Coulomb fields. Show that if  $q_1/m_1 = q_2/m_2$ , then the dipole radiation vanishes.

### Problem 3 (25 points):

The magnetic dipole field has exactly the same form as the electric dipole field, except that the electric dipole  $\vec{d}$  is replaced by the magnetic dipole  $\vec{M}$ . Here we consider a spinning neutron star that is slowing down due to magnetic dipole radiation. Let us assume that the neutron star is a uniform sphere, with mass  $M$ , radius  $R$ , spinning at angular velocity  $\omega$ . Let us also assume that its magnetic field is a dipole, with a strength  $B$  at the magnetic pole. In general, the angle between the magnetic and rotation axes is  $\theta$ , i.e. the dipole is rotating.

(a) Show that the magnitude of the magnetic dipole moment  $M$ , is  $M = BR^3/2$ , and find an expression for the power radiated by the magnetic dipole, in terms of  $B$ ,  $R$ ,  $\omega$ , and  $\theta$ . (Hint: compute the E field of a dipole – two charges separated by a small distance.)

(b) Find an expression for the time it takes for the neutron star to lose all of its rotational energy. What is this “spin-down” time for a typical neutron star (assume  $M = 1M_\odot$ ,  $R = 10^6$  cm,  $B = 10^{12}$  gauss,  $\omega = 10^3$  s, and  $\theta = \pi/2$ ).

### Problem 4 (25 points):

Consider the semi-classical Bohr model of the hydrogen atom, with the electron in circular orbit around a proton, radiating continuously according to the Larmor formula.

(a) Starting from an orbit with the energy equal to the  $n = 2$  level (i.e. the first excited quantum state), how long would it take for the electron to descend to the ground state?

(b) Compare this with the actual quantum-mechanical lifetime (hint: you will have to look up the Einstein coefficient  $A_{21}$  for the Lyman  $\alpha$  transition).