

# Radio Recombination Lines from Rydberg atoms

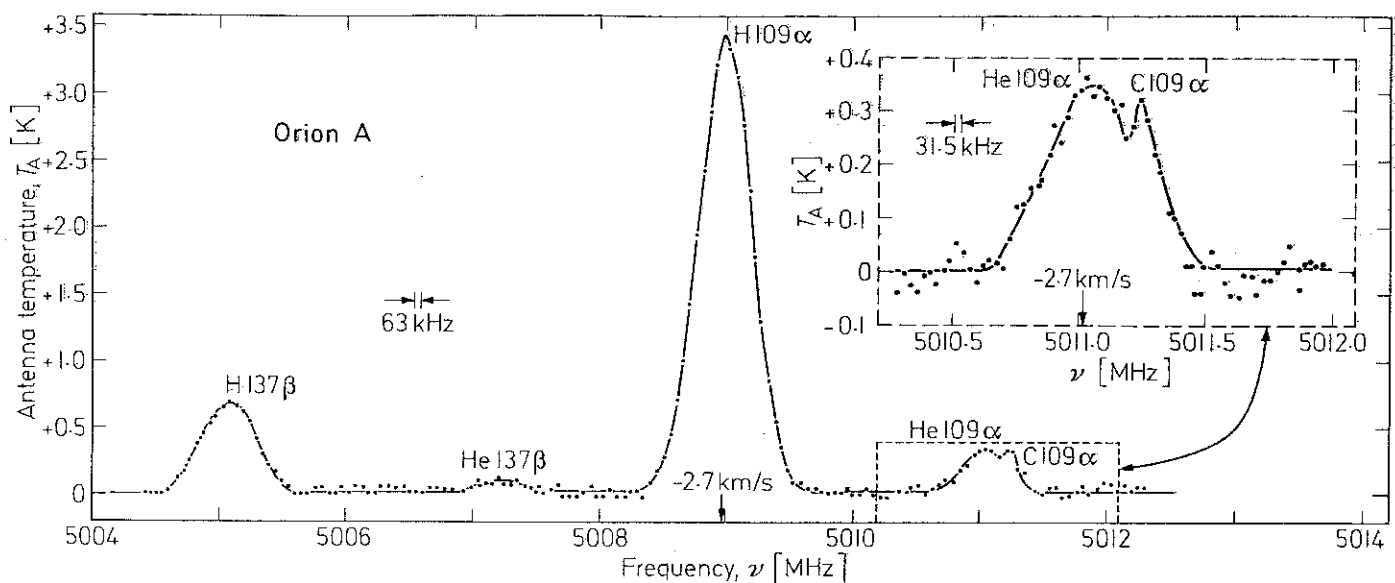


Fig. 12. Broadband spectrogram of the  $109\alpha$  region of the spectrum of the Orion Nebula. The frequency resolution is 63 kHz for the broadband spectrogram and 31.5 kHz for the narrow band spectrum centered on the He  $109\alpha$  line. (After CHURCHWELL and MEZGER, 1970, by permission of Gordon & Breach Science Publishers)

Hydrogenic atoms' wavelengths

$$\frac{1}{\lambda} = \frac{\mu e^4}{4\pi \hbar^3 c} \left( \frac{1}{n^2} - \frac{1}{m^2} \right)$$

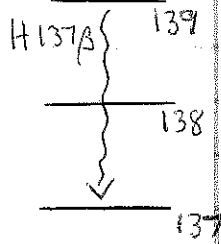
$n$  = lower level  
 $m$  = upper level  
 $\mu$  = reduced mass

where  $\mu = \frac{m_e m_p}{m_e + m_p}$  for hydrogen

$\mu = m_e$  for infinite mass nucleus

Rydberg constant

$$R_H = \frac{\mu e^4}{4\pi \hbar^3 c} = 109677.60 \text{ cm}^{-1} = \frac{1}{912 \text{ \AA}}$$



Rydberg constant  $R_\infty = \frac{m_e e^4}{4\pi \hbar^3 c} = 109737.31 \text{ cm}^{-1}$

For example,  $n=109$ ,  $m=110$  line is called H  $109\alpha$

$$\nu = c/\lambda = 5008.92 \text{ MHz for hydrogen } (\lambda = 5.985 \text{ cm})$$

$$\nu = c/\lambda = 5011.65 \text{ MHz for infinite mass nucleus}$$

