

Figure 34.1 Radiative cooling function $\Lambda/n_e n_H$ for solar-abundance plasma in CIE, computed with the CHIANTI code (Dere et al. 2009). Dashed lines show simple power-law approximations for $10^5 < T < 10^{7.3}$ K and for $T > 10^{7.5}$ K.

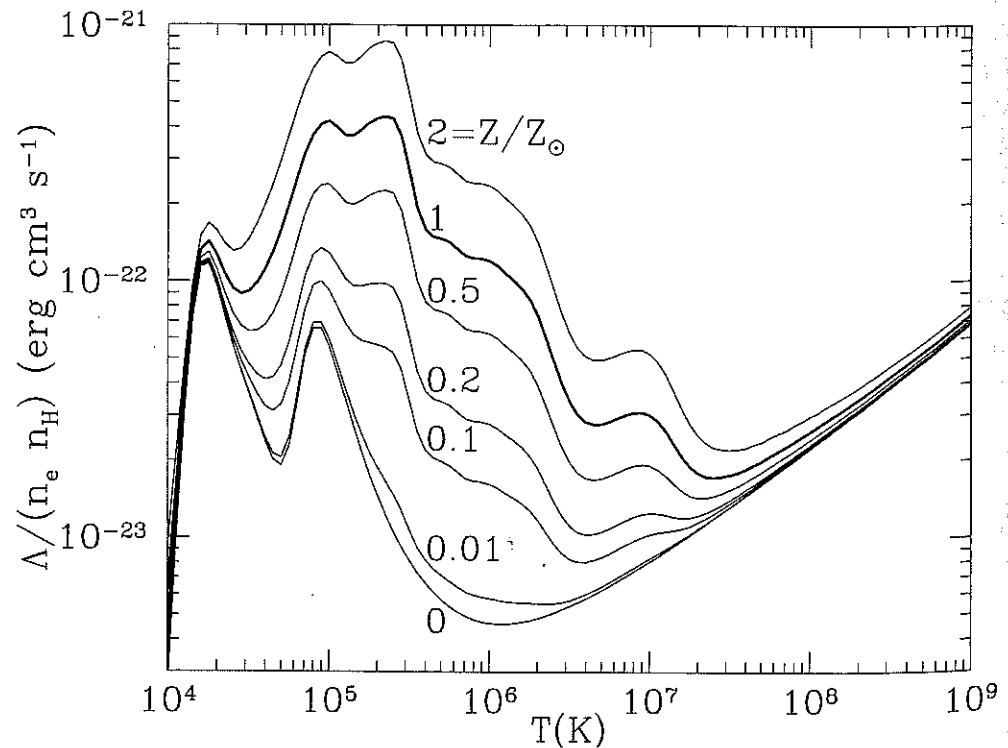


Figure 34.2 Same as Fig. 34.1, but for different abundances. Z/Z_\odot is the abundance of elements heavier than He relative to solar abundances.